

Concludes Calculation of Fluctuation Spectrum.

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ii.) Test Particle Model - Fluctuation Spectrum

→ Basic ideas:

- thermal equilibrium of (stable) plasma is balance of:

→ Cerenkov emission - of plasma waves by discrete particles (i.e. wake)

→ absorption of waves by Landau damping.

- Key ideas:

→ weak fluctuations - linear trajectories (unperturbed orbits)

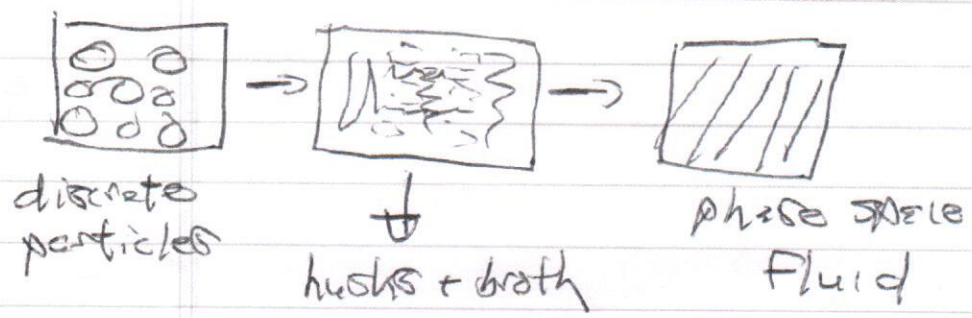
→ uncorrelated emission (of) test particles

→ each particle a "double agent":

1) → discrete emitter

2) → part of the phase space fluid which absorbs the emission from other test particles.

n.b.: $N_{CSMG} \sim$ pea soup (husks + soup)



"every pea in the pea soup is part of the soup to other peas"

Consider electron N_{CSMG} , with stationary cond

For test particle potential, extend Debye calculation:

$$dF = F_C + \tilde{F}$$

\downarrow \downarrow
 coherent discreteness
 v/c < v source
 response

$$dF = \frac{e^2 E_{k_0}^2}{m \cdot (-i(\omega - kv))} \frac{\partial K_F / \partial v}{\partial(x-x_H)} \partial(v-v_H)$$

\downarrow \downarrow
 coherent response discrete source

$$\begin{aligned}\nabla^2 \phi &= 4\pi n_0 e |\int dV df \\ &= 4\pi n_0 e |\int dV f^0 + 4\pi n_0 e |\int dV \tilde{f}\end{aligned}$$

so

$$E(k, \omega) \hat{\phi}_{k, \omega} = \frac{4\pi n_0 e}{k^2} \int dV \tilde{f}_{k, \omega}$$

Now, using u.p.o.:

$$\int \tilde{f}_k = \int dx e^{-ikx} |\epsilon| \delta(x - x(t))$$

$$x(t) = x_0 + vt$$

so

$$\hat{\phi}(k, t) = \frac{4\pi n_0 e}{k^2} e^{-ikvt}$$

\Rightarrow driven solution (discretized u.p. damping)

$$\hat{\phi}_k(t) = \frac{4\pi n_0 e}{k^2} e^{-ikvt}$$

$$+ \hat{\phi}_k^{\text{homos}} e^{-i\omega t}$$

homos. solution

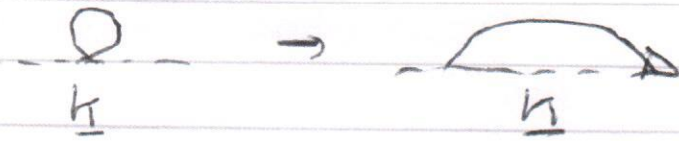
$$\omega_y = \omega_n(k) + i\omega_d(k)$$

so time asymptotically:

- $\omega_d(k) < 0 \Rightarrow$ homogeneous collective response damped.
- only driven (by discreteness) solutions persist

but:

- $\omega_d \lesssim 0$, may need wait long time ^{fluctuations in}
- for sufficient source strength, system may grow to nonlinearity before damping occurs.
- if unstable modes require ultimate nonlinear damping to balance noise
 i.e. $E_{IM} = E_{IM}(k, \omega, \langle \hat{\phi}^2 \rangle)$
 \rightarrow "noise" = thermal + nonlinear, then
 \rightarrow transfer, as well as emission and absorption, occurs.

then 

→ Have:

$$\langle \hat{\phi}^2 \rangle_{k, \omega} = \left(\frac{4\pi m_0 k e^2}{k^2} \right)^2 \int dV_1 \int dV_2 \frac{\langle \tilde{f}(V_1) \tilde{f}(V_2) \rangle_{k, \omega}}{|\epsilon(k, \omega)|^2}$$

∴ all constant:

- Coulomb factor

- discreteness source $\langle \tilde{f} \tilde{f} \rangle_{k, \omega}$

- $\epsilon(k, \omega)$ collective response

→ Noise

$$\tilde{f} = \frac{q}{N} \sum_{i=1}^N \delta(x - x_i(t)) \delta(v - v_i(t))$$

→ u.p.o. $\begin{cases} x_i(t) = x_{i0} + v_i t \\ v_i(t) = v, \text{ const} \end{cases}$

$$\langle \rangle = n \int dx_i \int dv_i \langle F(v_i) \rangle$$

↳ Maxwellian

avg. over eqbm distrib. of
discr. uncorrelated test particles

"uncorrelated" → dilute, $k_B T \gg e^2 / r$
→ $\frac{1}{2} n \lambda_D^3 \ll 1$

JB

$$\langle \tilde{F}(\underline{u}) \tilde{F}(\underline{v}) \rangle = \left\langle n \int dx_i \int dv_i \left(\frac{1}{N} \sum_{\alpha=1}^N \delta(\underline{x}_i - \underline{x}_{\alpha}(t)) \delta(\underline{v}_i - \underline{v}_{\alpha}(t)) \right) \left(\frac{1}{N} \sum_{\beta=1}^N \delta(\underline{v}_j - \underline{v}_{\beta}(t)) \delta(\underline{x}_j - \underline{x}_{\beta}(t)) \right) \right\rangle$$

only ≠ 0 if arguments interchangeable!

$$= \int dx_i \int dv_i \frac{\langle F \rangle}{N} \sum_{i \neq j}^N \delta(\underline{x}_i - \underline{x}_j) \delta(\underline{x}_i - \underline{x}_j) \delta(\underline{v}_i - \underline{v}_j) \delta(\underline{v}_i - \underline{v}_j)$$

$$= \frac{\langle F \rangle}{N} \delta(\underline{x}_i - \underline{x}_j) \delta(\underline{v}_i - \underline{v}_j)$$

So, discreteness correlation function:

$$\langle \tilde{F}(\underline{1}) \tilde{F}(\underline{2}) \rangle = \frac{\langle F \rangle}{N} \delta(\underline{x}_1 - \underline{x}_2) \delta(\underline{v}_1 - \underline{v}_2)$$

- i.e.
- no width, no range
 - particles only correlated if same particle.

Now, need:

$$\langle \tilde{F}(\underline{1}) \tilde{F}(\underline{2}) \rangle_{k, \omega}$$

$$\langle \tilde{F}(\underline{1}) \tilde{F}(\underline{2}) \rangle_k = \int e^{-ik(x_2 - x_1)} \langle \tilde{F}(\underline{1}) \tilde{F}(\underline{2}) \rangle$$

and ~~the~~ time transform \leftrightarrow time history (u.p.o.)

$$\langle \tilde{F}(\underline{1}) \tilde{F}(\underline{2}) \rangle_{k, \omega} \equiv \int_0^{\infty} dt e^{i\omega t} u(\underline{2}, T) \int e^{-ik(x_2 - x_1)} \langle \tilde{F}(\underline{1}) \tilde{F}(\underline{2}) \rangle$$

$$+ \int_{-\infty}^0 dt e^{i\omega t} u(\underline{1}, -T) \int e^{-ik(x_2 - x_1)} \langle \tilde{F}(\underline{1}) \tilde{F}(\underline{2}) \rangle dx$$

$U \rightarrow$ operator pushing particle along
u.p.o.

$$U: X \rightarrow X + VT$$

$$\textcircled{2}: X_2 \rightarrow X_2 + V_2 T$$

$$\textcircled{2} = \int_0^{\infty} dt \int_0^{\infty} dx_2 e^{-ik(x_2 - x)} e^{i(\omega - kv_2)t} \langle \tilde{F}(t) \tilde{F}(x) \rangle dx_2$$

$$= \int_0^{\infty} dt e^{i(\omega - kv_2)t} \langle \tilde{F}(t) \tilde{F}(x) \rangle_k$$

$$= \frac{-1}{i(\omega - kv_2)} \langle \tilde{F} \tilde{F} \rangle_k$$

$$= \frac{i}{(\omega - kv_2)} \langle \tilde{F} \tilde{F} \rangle_k$$

seek $\frac{re}{-}$

$$= \pi \delta(\omega - kv_2) \langle \tilde{F} \tilde{F}(k) \rangle$$

Similarly,

$$\textcircled{1} = \pi \delta(\omega - kv_1) \langle \tilde{F} \tilde{F} \rangle_k$$

~~and~~

and seek:

$$\int dv_1 \int dv_2 \langle \tilde{F}(v_1) \tilde{F}(v_2) \rangle_{v_1, v_2}$$

$$\overline{\langle \tilde{F}(t) \tilde{F}(-) \rangle} = \overline{\langle F \rangle} \delta(x) \delta(v_-)$$

$$\langle \tilde{F} \tilde{f} \rangle_{\kappa} = \overline{\langle F \rangle} \delta(v_-)$$

$$\int dv_1 \int dv_2 = \int dv_+ \int dv_-$$

$v_- \rightarrow$ relative

$v_+ \rightarrow$ CM

$$\int dv_- \langle \tilde{F}(t) \tilde{F}(z) \rangle_{\kappa, \omega} = 2\pi \delta(\omega - \kappa v_-) \overline{\langle F \rangle}$$

and

$$\left\langle \frac{\hat{n}}{n_0} \frac{\hat{n}}{n_0} \right\rangle_{\kappa, \omega} = \int dv_- \overline{\langle F \rangle} 2\pi \delta(\omega - \kappa v_-)$$

$$\equiv c(\kappa, \omega)$$

\downarrow
emission correlator.

\rightarrow v_{the} extracted.

$$C(k, \omega) = \frac{2\pi}{n|k|v_{the}} \langle \tilde{F}(\omega/kv_{the}) \rangle$$

$$\downarrow = \frac{2\pi}{n} \tilde{F}_{str} \langle \tilde{F}(\omega/kv_{the}) \rangle$$

\rightarrow streaming time

discreteness noise has Maxwellian Doppler spectrum (obvious)

so, thermal equilibrium spectrum:

$$\langle \tilde{\phi}^2 \rangle_{k, \omega} = \left(\frac{4\pi n_0 e^2}{k^2} \right)^2 \frac{1}{n_0 |k| v_{the}} \frac{2\pi}{|\epsilon(k, \omega)|^2} \langle \tilde{F}(\omega/kv_{the}) \rangle$$

so, spectrum set by:

- equilibrium particle emission distribution $\sim \omega^2/k^2 v_{the}^2$

- collective resonances, i.e.

$$\omega \geq kv_{the} \quad \epsilon \approx 1 - \omega_p^2/\omega^2 + i\epsilon IM$$

$$\omega < kv_{th} \quad \epsilon \approx 1 + 1/k^2 \lambda_D^2 + i\epsilon IM$$

- Coulomb Factor (screening modified)
- T_{otr} (from time transform)

00

- collective response strongest at wave resonance ($kV \sim \omega_p \sim \omega$)
- ⇒ expect peak in frequency spectrum

(n.b. emission distribution hits wave resonance $kV_{th} \sim \omega_p \Rightarrow k \sim \lambda^{-1}$,
for scale)

Limiting behavior:

- For $\omega > \omega_p$, noise source decouples

from collective dynamics (i.e. $\epsilon \rightarrow 1$ as $\omega \gg \omega_p$);

$$\langle \phi^2 \rangle_{k,\omega} \approx N_0 \left(\frac{4\pi k l}{k^2} \right)^2 \frac{2\pi}{k l v_{the}} e^{-\omega^2 / k^2 v_{the}^2}$$

- for $\omega < \omega_p$, low frequency noise \ominus static \Rightarrow screened by $n_0 e_0 m q$

∞

$$\langle \vec{p}^2 \rangle_{k, \omega} \approx N_0 \frac{(4\pi k e)^2}{|k| v_{th}} \frac{2\pi e^{-\omega^2/k^2 v_{th}^2}}{\left(k^2 + 1/\lambda_D^2\right)^2}$$

\uparrow
 abs k^2

∞

can write Electric Field Spectrum

$$\frac{\langle F \rangle}{|k| v_{th}} \sqrt{2\pi}$$

\uparrow
 \rightarrow re-describes

$$\frac{\langle E^2 \rangle_{k, \omega}}{8\pi} = \frac{4\pi^2 N_0 e^2}{k^2 |k|} \frac{F(\omega/k v_{th})}{\left[\left(1 - \frac{\omega_p^2}{\omega^2}\right)^2 + \left(\frac{\pi \omega_p F'}{|k| k}\right)^2 \right]}$$

$$\left(F' = \frac{dF}{du} \right) \quad |E_{rI}|^2 + |E_{IM}|^2$$

$u = \omega/k$

i.e. Thermal E-field Spectrum

$$\frac{\langle E^2 \rangle_{k, \omega}}{8\pi} = \frac{4\pi^2 N_0 e^2}{k^2 |k|} \left(F / |E_{rI}|^2 + |E_{IM}|^2 \right)$$

Now to make contact with usual expectations of "k_BT/2 per d.o.f".

$$W_n = \int \frac{d\omega}{2\pi} \langle E^2 \rangle_{n,\omega} / 2\pi$$

↓
field energy
per mode

Useful trick; Pole Approximation:

$$1/|E|^2 = \frac{1}{\left[(\omega - \omega_n)^2 \left| \frac{\partial E}{\partial \omega} \right|^2 + |E_{IM}|^2 \right]}$$

↓
real frequency
sets location

↓
width

$$\approx \frac{1}{|E_{IM}|} \left[\frac{|E_{IM}|}{(\omega - \omega_n)^2 \left| \frac{\partial E}{\partial \omega} \right|^2 + |E_{IM}|^2} \right]$$

$$\approx \frac{1}{|E_{IM}|} \left| \frac{\partial E}{\partial \omega} \right|_{\omega_n}^{-1} \pi \delta(\omega - \omega_n)$$

↓
exists on collective
resonance

so, pole approximation:

$$\frac{1}{|E|^2} = \frac{\pi \delta(\omega - \omega_n)}{|E_{\text{Im}}(\omega_n)| \left| \frac{\partial E_r}{\partial \omega} \right|_{\omega_n}}$$

so integrating in pole approx:

$$\begin{aligned} W_n &= \frac{m_0 \omega_p^2}{2k} \frac{F}{|F'|} \\ &= m_0 \frac{\omega_p^2}{2k} \frac{F}{\frac{\omega_p}{|k|} \frac{F}{v_c^2}} = T/2 \end{aligned}$$

in accord with "T/2 per d.o.f" intuition,

$$\rightarrow \text{if } k \lambda_D \gg 1, \quad E \approx 1 + 1/k^2 \lambda_D^2$$

no collective resonance

$$W_n \approx \frac{T}{2} \frac{1}{k^2 \lambda_D^2} \rightarrow \text{strong cut-off beyond } \lambda_D.$$

so, for total energy density: (30)

$$\langle E^2/\epsilon_{II} \rangle = \int d\mathbf{k} \omega_{\mathbf{k}}$$

$$\sim \left(\frac{k_B T}{2} \right) k_{\max}^3$$

$$\sim n \frac{k_B T}{2} \frac{1}{n \lambda_D^3}$$

$$\sim (P_{KE}) / n \lambda_D^3$$

↓
in Debye sphere

consistent with idea of diluteness:

$$(FED) \sim (P_{KE}) / n \lambda_D^3$$

$1/n \lambda_D^3 \sim$ diluteness/discreteness factor

→ To connect formally, to Fluctuation-dissipation theorem:

Notes:
$$\epsilon_{IM} = \frac{-\omega_p^2 \pi}{k |k|} \frac{\partial \langle f \rangle}{\partial V} \Big|_{\omega/k}$$

$$= \frac{2\pi \omega}{k^2 v_{te}^2} \frac{\omega_p^2}{|k| v_{te}} \langle \bar{f}(\omega/k) \rangle$$

for Maxwellian $\langle f \rangle$

so $\langle \bar{f}(\omega/k) \rangle = k^2 v_{te}^2 / |k| v_{te} \epsilon_{IM} / 2\pi \omega \omega_p^2$

we have:

$$\langle \hat{\phi}^2 \rangle_{k,\omega} = \frac{2\pi T}{|k| v_{te}} \left(\frac{4\pi |e|}{k^2} \right)^2 \frac{\langle \bar{f}(\omega/k) \rangle}{|E(k,\omega)|^2}$$

so, plugging in:

$$\langle \hat{\phi}^2 \rangle_{k,\omega} = \frac{8\pi T}{k^2 \omega} \frac{\text{Im} \epsilon}{|\epsilon|^2}$$

and

$$\left\langle \frac{\hat{E}^2}{8\pi} \right\rangle_{k,\omega} = \frac{T}{\omega} \frac{\text{Im} \epsilon}{|\epsilon|^2}$$

$\langle \hat{\phi}^2 \rangle \sim \frac{1}{\omega} \frac{\text{Im} \epsilon}{|\epsilon|^2}$

Fluctuation-Dissipation Theorem

(restated form of spectrum)

→ relates thermal fluctuations to dissipation in collective modes (Im ϵ)

→ obviously consistent (by construction), with physical picture.

i) Some general comments:

→ key element of T.P.M. is $\left\{ \begin{array}{l} \text{causality} \\ \text{use of} \\ \text{linear } F_{k,u}^c \text{ or, equivalently, unperturbed orbit} \end{array} \right.$

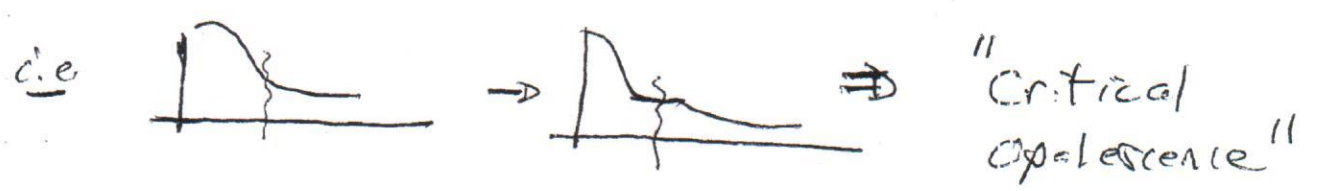
- This assumes small fluctuation levels, so stochastic deflection is 'weak'

d.e. $\underline{x}(t) = \underline{x}(0) + \underline{v}t + \int_0^t \underline{d}(\underline{x}(t))$
deflection

How weak? \leftrightarrow take care $\gamma_{ec} < \gamma_T$ condition

but - $\langle \hat{\phi}^2 \rangle_k \sim () \frac{F(\omega_{pe}/k)}{|F'(\omega_{pe}/k)|}$

Fluctuations diverge as $F' \rightarrow 0$, from below



Note $F' > 0$ not necessary \leftrightarrow theory

fails for stable plasma -----, approaching marginality.



→ As fluctuations grow, linearizations fail

∴ must renormalize!

→ particle propagator

$$c / \omega - kv \rightarrow c / \omega - kv + \Sigma$$

Σ
self energy,
dissipation rate

→ mode propagator / response

$$I/E \rightarrow I \left[\underbrace{\omega - (\omega_k + \delta\omega_k)}_{\text{nonlinear frequency shift}} \frac{\partial \epsilon}{\partial \omega} + i \left(\underbrace{\epsilon_{IH}^L + \epsilon_{IH}^{NL}}_{\text{nonlinear dissipation}} \right) \right]$$

nonlinear
frequency
shift

nonlinear
dissipation
($\omega \rightarrow \omega$ interaction)
($\omega \rightarrow 0$ interaction)
⇒ γ_{NL}

(recall NL oscillator, driven)

Calculating all this is aim of plasma turbulence theory